



## CECAM Flagship School, October 2, 2023

# First steps with SIESTA: from zero to hero

# Basis sets in SIESTA

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# siesta



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First steps with SIESTA: from zero to hero

## Basis Expansion: ← M basis functions $|\psi_i\rangle = \sum c_{i,\mu} |\mu\rangle$ $\mu = 1$

Generalised eigenvalue problem:

 $H_{\mu\nu}C_{i\nu} = \varepsilon_i S_{\mu\nu}C_{i\nu}$ 

(matrix form)

 $H_{\mu\nu} = \langle \mu | \hat{H} | \nu \rangle \qquad S_{\mu\nu} =$ Hamiltonian matrix

 $S_{\mu\nu} = \langle \mu \mid \nu \rangle$ 

Overlap matrix

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Density Matrix





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# **Basis Expansion:** $|\psi_i\rangle = \sum c_{i,\mu} |\mu\rangle$ μ=1 $\langle r | \mu \rangle = \phi_{\mu}(r)$

 $n(r) = \sum_{i=1}^{N} |\psi_{i}(r)|^{2} = \sum_{i=1}^{N} \sum_{i=1}^{M} c_{\mu,i} c_{\nu,i}^{*} \phi_{\mu}(r) \phi_{\nu}^{*}(r) =$ 



Density Functional Theory (DFT) Background Periodic Boundary Conditions?: Bloch theorem  $|\psi_{i}\rangle = \sum_{\mu=1}^{M} c_{i,\mu}(\mu)$ Bloch orbitals  $\longrightarrow |n\vec{k}\rangle = \frac{1}{\sqrt{N}} \sum_{\mu,\vec{k}} c_{n,\mu}(\vec{k}) \cdot e^{i\vec{k}\cdot\vec{R}} |\mu\vec{R}\rangle$  $|\mu\vec{k}\rangle = \frac{1}{\sqrt{N}} \sum_{\vec{z}} e^{i\vec{k}\cdot\vec{R}} |\mu\vec{R}\rangle$ Generalised eigenvalue problem:

 $H_{\mu\nu}(\vec{k})C_{j\nu}(\vec{k}) = \varepsilon_{j,\vec{k}} S_{\mu\nu}(\vec{k})C_{j\nu}(\vec{k})$ 



## $H_{\mu\nu}(\vec{k}) = \langle \mu, \vec{k} | \hat{H} | \nu, \vec{k} \rangle = \sum e^{i\vec{k}\cdot\vec{R}} \langle \mu, 0 | \hat{H} | \nu, \vec{R} \rangle$ $\overrightarrow{R}$



# What type of Basis Functions?

## Representation of Hilbert space for the electronic wavefunctions



### Planewaves

- Good for PBC
- Systematic convergence (complete)
- Orthogonal and Spatially unbiased
- Expressions for H are simple (FFT)

- Wavelets,
- Bessel functions,
- Psinc's, blips,
- Augmented Plane Waves,
- Muffin-tin Orbitals, etc.



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## Atomic Orbitals

Other Numerical Functions:

- Vacuum costs as much as matter!
- Large number of PWs per electron
- Compact orbitals are harder
- Localisation ideas are not easily implemented

# Atomic Orbitals

 $\phi_{\mu}(r - R_{\mu}) \rightarrow \phi_{Ilmn}(r_I) = R_{Iln}(r)Y_{lm}(\hat{r}_I)$ 

Atom index 🦯 angular momentum

Multiple orbitals with same lm

magnetic quantum number





 $\langle r | \mu \rangle = \phi_{\mu}(r) \longrightarrow \phi_{\mu}(r - R_{\mu})$ 

# LCAO methods

Orbitals close to the solutions of the atomic problem are good approximations

Tradition of Quantum Chemistry methods:

- Slater-type orbitals (STO)
- Gaussian-type orbitals (GTO)
- Contracted Gaussian-type orbitals
- Numerical real-space grid



# Atomic Orbitals

 $\phi_{\mu}(r - R_{\mu}) \rightarrow \phi_{Ilmn}(r_I) = R_{Iln}(r)Y_{lm}(\hat{r}_I)$ Atom index Multiple orbitals with same lm angular momentum magnetic quantum number dyz  $d_{x^{2}-y^{2}}$  $d_{xz}$ d<sub>z</sub>2 d <sub>xy</sub>

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 $\langle r | \mu \rangle = \phi_{\mu}(r) \longrightarrow \phi_{\mu}(r - R_{\mu})$ 

## LCAO methods

- Very efficient in terms of number of orbitals per electrons
- Very well suited to describe localisation
- Large reduction in CPU and memory costs
- No need for periodicity
- Vacuum is almost free
- Chemical information (charge population, projected density of states, etc).
- Lack of systematics for convergence.
- Require human and computational effort to get a good basis set before use Spatially biased, since they are optimal for an atomic problem: Basis Set Superposition Error
  - Orbitals move with atoms, which brings extra terms in forces (Pulay corrections)
- Calculation of Hamiltonian matrix elements can be quite complicated (and expensive)

# Atomic Orbitals

# $\phi_{\mu}(r-R_{\mu}) \rightarrow \phi_{Ilmn}(r_{I}) = \left(R_{Iln}(r)Y_{lm}(\hat{r}_{I})\right)$ Radial flexibility

## $\odot$ Minimal basis, a.k.a. single- $\zeta$ (SZ)

 One single radial function per angular momentum. shell occupied in the free atom

## $\odot$ Multiple- $\zeta$

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- Add more than one radial function with the same. angular momentum.
- Double- $\zeta$  (DZ), Triple- $\zeta$  (TZ), cuadruple- $\zeta$  (QZ), etc.

## Diffuse functions

 Add radial function on pre-existing valence shell, with longer tail than free atom

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 $\langle r | \mu \rangle = \phi_{\mu}(r) \longrightarrow \phi_{\mu}(r - R_{\mu})$ 









# Hierarchy of basis sets Standardised tiers

- General procedure:
  - Start from SZ
  - Every step: increase both  $\zeta$  and polarisation by one

Quick/cheap exploratory calculations

single- $\zeta$ 



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General philosophy

The larger the number of orbitals, the better the quality

Highly converged calculations

+ multiple-ζ
+ polarisation
+ diffuse orbitals





# Hierarchy of basis sets

# Standardised tiers

Atom	Valence	SZ		DZ		Р	
	configuration						
		# orbita	ls symmetry	# orbita	ls symmetry	# orbitals	symmetry
Si	$3s^2 \ 3p^2$	1	S	2	S	1	$d_{xy}$
		1	$p_x$	2	$p_x$	1	$d_{yz}$
		1	$p_y$	2	$p_y$	1	$d_{zx}$
		1	$p_z$	2	$p_z$	1	$d_{x^2-y^2}$
						1	$d_{3z^2-r^2}$
	Total	4		8		(DZ+P) 13	

Atom	Valence						
	configuration						
		# orbita	${ m lssymmetry}$	# orbitals	s symmetry	# orbitals	symmetry
Fe	$4s^2 \ 3d^6$	1	s	2	S	1	$p_x$
		1	$d_{xy}$	2	$d_{xy}$	1	$p_y$
		1	$d_{yz}$	2	$d_{yz}$	1	$p_{ m z}$
		1	$d_{zx}$	2	$d_{zx}$		
		1	$d_{x^2-y^2}$	2	$d_{x^2-y^2}$		
		1	$d_{3z^2-r^2}$	2	$d_{3z^2-r^2}$		
	Total	6		12		(DZ+P) 15	



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# Hierarchy of basis sets Convergence with size of basis set

- Polarisation orbitals are important for convergence.
- DZP basis is usually "good enough" as compared to converged PW calculations.
- Deviations are comparable to differences due to pseudopotentials or XC functionals

TABLE II. Basis comparisons for bulk Si.  $a, B, and E_c$ stand for lattice parameter (in Å), bulk modulus (in GPa), and cohesive energy (in eV), respectively. SZ, DZ, and TZ stand for single- $\zeta$ , double- $\zeta$ , and triple- $\zeta$ . P stands for polarized, DP for doubly polarized. LAPW results were taken from Ref. 48, and the experimental values from Ref. 39.

	SZ	DZ	TZ	SZP	DZP	TZP	TZDP	$\mathbf{PW}$	LAPW	Exp
$\overline{a}$	5.52	5.49	5.48	5.43	5.40	5.39	5.39	5.38	5.41	5.43
B	85	87	85	97	97	97	97	96	96	98.8
$E_c$	4.70	4.83	4.85	5.21	5.31	5.32	5.34	5.40	5.28	4.63





Junquera et al Phys. Rev. B 64, 235111 (2001)



# Atomic orbitals in SIESTA

General functional form: 

$$\phi_{Ilmn}(r_I)$$
 =

- > Strictly localised (zero beyond a certain cutoff radii,  $r_c$ )
  - $\checkmark$  As many as you want (I-shells and number of  $\zeta$ s)
  - ✓ Of any (radial) shape
  - $\checkmark$  Any localisation range (r<sub>c</sub>)
  - $\checkmark$  Any center (not necessarily an atom)
  - Numerical Pseudo-Atomic Orbitals (PAO)



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## $= R_{Iln}(r)Y_{lm}(\hat{r}_{I})$







# Atomic orbitals in SIESTA

## Radial functions can be:

- Introduced by user (deprecated)
  - Basis type "user", read a file with a discretised table of values for r

## Generated by SIESTA:

- Numerical solution of pseudoatom + "modifications"
- Depends on parameters that need to be defined by user
- Quite tunable
- Various levels of automatism, and predefinition of default values for parameters

First steps with SIESTA: from zero to hero













- Solution of Kohn-Sham DFT on pseudo-atom under an added confinement potential
- Control the range of the orbitals with one parameter:

$$\left(-\frac{1}{2r}\frac{d^2}{dr^2}r + \frac{l(l+1)}{2r^2} + V_l(r)\right)R_l(r) = (\varepsilon_l + \delta\varepsilon_l)R_l(r)$$



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► FIREBALLS: O. F. Sankey & D. J. Niklewski, Phys. Rev. B 40, 3979 (1989)

Energy Shift

(Radial equation)

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Bound states in a "potential well" of arbitrary shape Complement MIII – "Quantum Mechanics", C. Cohen-Tannoudji et al.

E is quantised and only certain values are allowed

E > eigenenergy results in wfn  $\rightarrow -\infty$ 

 $\psi(r_c) = 0$  for certain point close to the well boundary





- Solution of Kohn-Sham DFT on pseudo-atom under an added confinement potential
- Control the range of the orbitals with one parameter:

$$\left(-\frac{1}{2r}\frac{d^2}{dr^2}r + \frac{l(l+1)}{2r^2} + V_l(r)\right)R_l(r) = (\varepsilon_l + \delta\varepsilon_l)R_l(r)$$

## The larger the Energy Shift, the shorter the $r_c$

Artacho et al Phys. Stat. Solidi (b) 215, 809 (1999)

-Cecan

First steps with SIESTA: from zero to hero

► FIREBALLS: O. F. Sankey & D. J. Niklewski, Phys. Rev. B 40, 3979 (1989)

Energy Shift

(Radial equation)











VS

WARNING: Basis can be optimised to get better agreement with experiments. That doesn't necessarily mean the basis set is better. There are other approximations implied (e.g. the xc functional) that could be responsible for disagreements between calculations and experiments!!

Bulk silicon – same cutoff radii for s & p orbitals

Soler et al J. Phys.: Condens. Matter, 14, 2745 (2002)

First steps with SIESTA: from zero to hero

- Convergence with the cutoff radii
- Larger radii  $\longrightarrow$  More accurate

## Shorter radii $\longrightarrow$ More efficient (computationally)









# Strictly localised orbitals? Soft-Confinement potentials Default in SIESTA is a hard-wall potential V(r < a) = 0 $V(r \ge a) = \infty$ Orbitals with discontinuous first derivative at $r_c$

Other functional forms are possible.

$$V(r) = V_0 \frac{e^{-\frac{r_c - r_i}{r - r_i}}}{r_c - r}$$

- + stricly localised (zero at  $r_c$ )
- + two parameters to optimise



## + orbitals with continuos derivatives



# Expanding radial flexibility Schemes to generate multiple- $\zeta$ basis sets



Ozaki & Kino Phys. Rev. B 69, 195113 (2004)



First steps with SIESTA: from zero to hero

Pseudoatom-wavefunctions with increasing number of nodes:

- + Orthogonality
- + Asymptotically complete (within sphere)
- unbound excited states of pseudos
- Efficiency requires longer rc







# Expanding radial flexibility Schemes to generate multiple- $\zeta$ basis sets





# "Split-valence" method

- Choose r<sub>m</sub> and continue smoothly towards the origin as  $r^l(a_l - b_l r^2)$
- Two parameters: the new orbital and its first derivative must be continuous at rm
- The second- $\zeta$  is the (normalised) difference between the first- $\zeta$  and the function above.





# Expanding radial flexibility Schemes to generate multiple- $\zeta$ basis sets





First steps with SIESTA: from zero to hero

# 6.0

# "Split-valence" method

- Choose r<sub>m</sub> and continue smoothly towards the origin as  $r^l(a_l - b_l r^2)$
- Two parameters: the new orbital and its first derivative must be continuous at rm
- The second- $\zeta$  is the (normalised) difference between the first- $\zeta$  and the function above.
- rm is controlled with PAO.SplitNorm (default is 0.15)

## Expanding angular flexibility Two schemes to generate polarization orbitals Atomic polarisation Perturbative polarisation

- Free pseudo atom valence orbitals under external electric field
- I+1 orbitals with same range of unperturbed orbitals Usually unbound: require short cutoffs



First steps with SIESTA: from zero to hero

Free pseudo atom orbitals of higher angular momentum

E. Artacho et al., Phys. Stat. Solidi (b) 215, 809 (1999)



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# Standardised tiers

## • General procedure:

- Start from SZ
- Every step: increase both  $\zeta$  and polarisation by one

Quick/cheap exploratory calculations

single- $\zeta$ 



First steps with SIESTA: from zero to hero

General philosophy

 The larger the number of orbitals, the better the quality

> Highly converged calculations

+ multiple-ζ
+ polarisation
+ diffuse orbitals







# Standardised tiers

### General procedure:

- Start from SZ
- Every step: increase both  $\zeta$  and polarisation by one

## SIESTA basis

- Multiple options to generate basis sets
- Hierarchical structure,
- Decent default basis...



First steps with SIESTA: from zero to hero

## General philosophy

The larger the number of orbitals, the better the quality

## BUT: optimal orbitals are environment-dependent!!





# How to input basis sets?

Test before going serious:

Although the default basis is usually a good starting point, it is important that you verify it is appropriate for your system.

If going to calculate for months or years, it is a good idea to spend a few days trying out bases and testing parameters is definitely worth it!





(C orbitals in diamond are not necessarily good for molecular CO2, or graphene)

Sometimes people can share bases in communities, the SIESTA mailing list, or third parties.

Practical session: Tuesday @12:30



# How to input basis sets?

	FDF flags	Default	Slightly better
Basis Size:	PAO.BasisSize	DZP	
Range of first-zeta:	PAO.EnergyShift	0.02 Ry	3-5 mRy
Second-zeta:	PAO.BasisType	Split	
Range of second-zeta:	PAO.SplitNorm	0.15	
Confinement:	PAO.SoftDefault	False	True
	PAO.SoftInnerRadius		0.9



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## PAO.SoftInnerRadius

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# How to input basis sets?

## Better control on the basis?

<pre>%block PAO.BasisSize</pre>	%block PAO.Basis				
Si DZ	O 2				
H SZP	n=2 0 2 E 50. 2.5				
O DZP	0.0 0.0				
<pre>%endblock PAO.BasisSize</pre>	1.0 1.0				
	n=2 1 2 P 1				
	0.0 0.0				
	1.0 1.0				
	<pre>%endblock PAO.Basis</pre>				





First steps with SIESTA: from zero to hero

%block PAO.Basis 0 3 n=2 0 2 0.0 0.0 1.0 1.0 n=2 1 2 0.0 0.0 1.0 1.0 n=3 2 1 0.0 1.0 %endblock PAO.Basis

Practical session: Tuesday @12:30













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Questions?

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![](_page_27_Picture_8.jpeg)